



University of  
New Hampshire



# Bridging instrument and science capabilities and performance

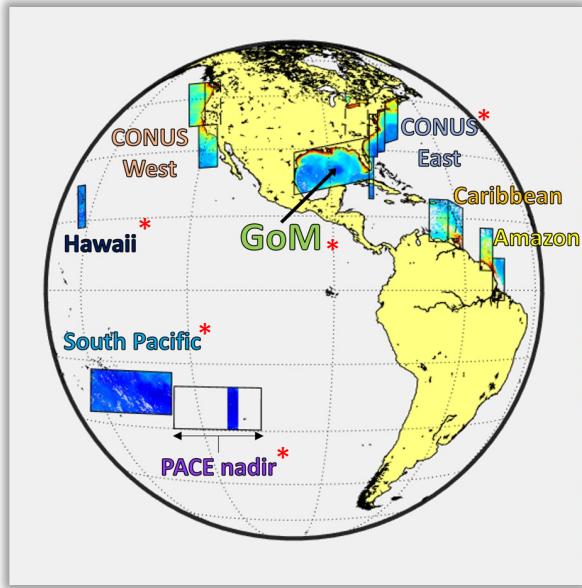
**IOCS 2023 SAT Meeting  
14 November**

*Antonio Mannino*

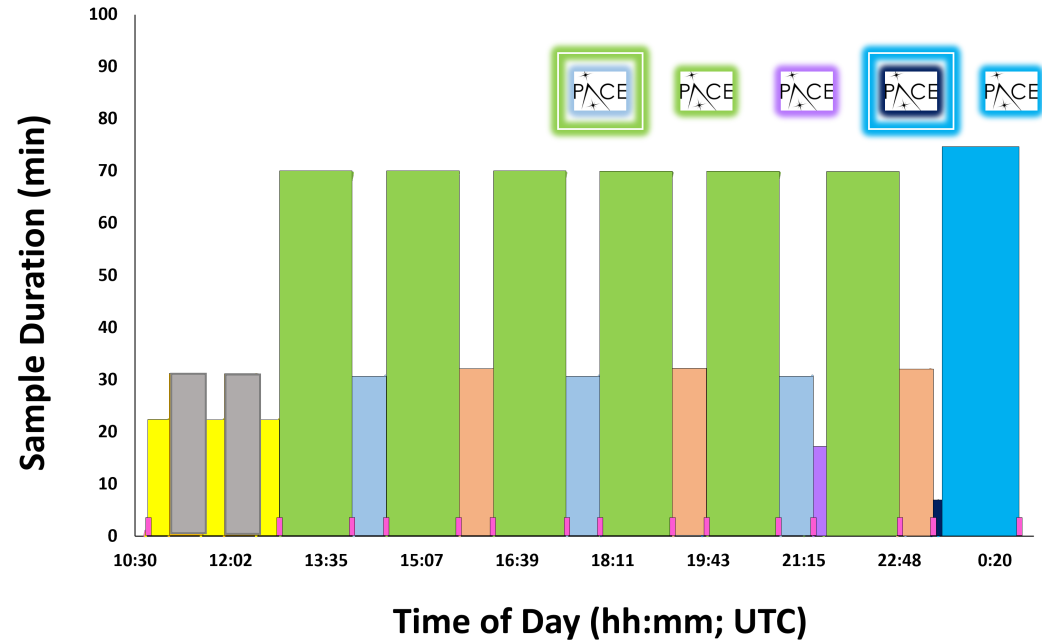
*Contributions from: Steve Persh, Ryan  
Vandermeulen, Sean Bailey & Boryana Efremova*

- Instrument capabilities and expected performance
  - Daily science operations and calibrations
  - Spectral
  - SNR
  - Radiometric uncertainties
- On-orbit calibration
- GLIMR Data Product Science Performance & Modeling ( $\rho_w$ )
- Validation of Science Data Products
- Peak into Cloud statistics
- Science Data Segment – Algorithm Tool

## Daily Fall/Spring schedule

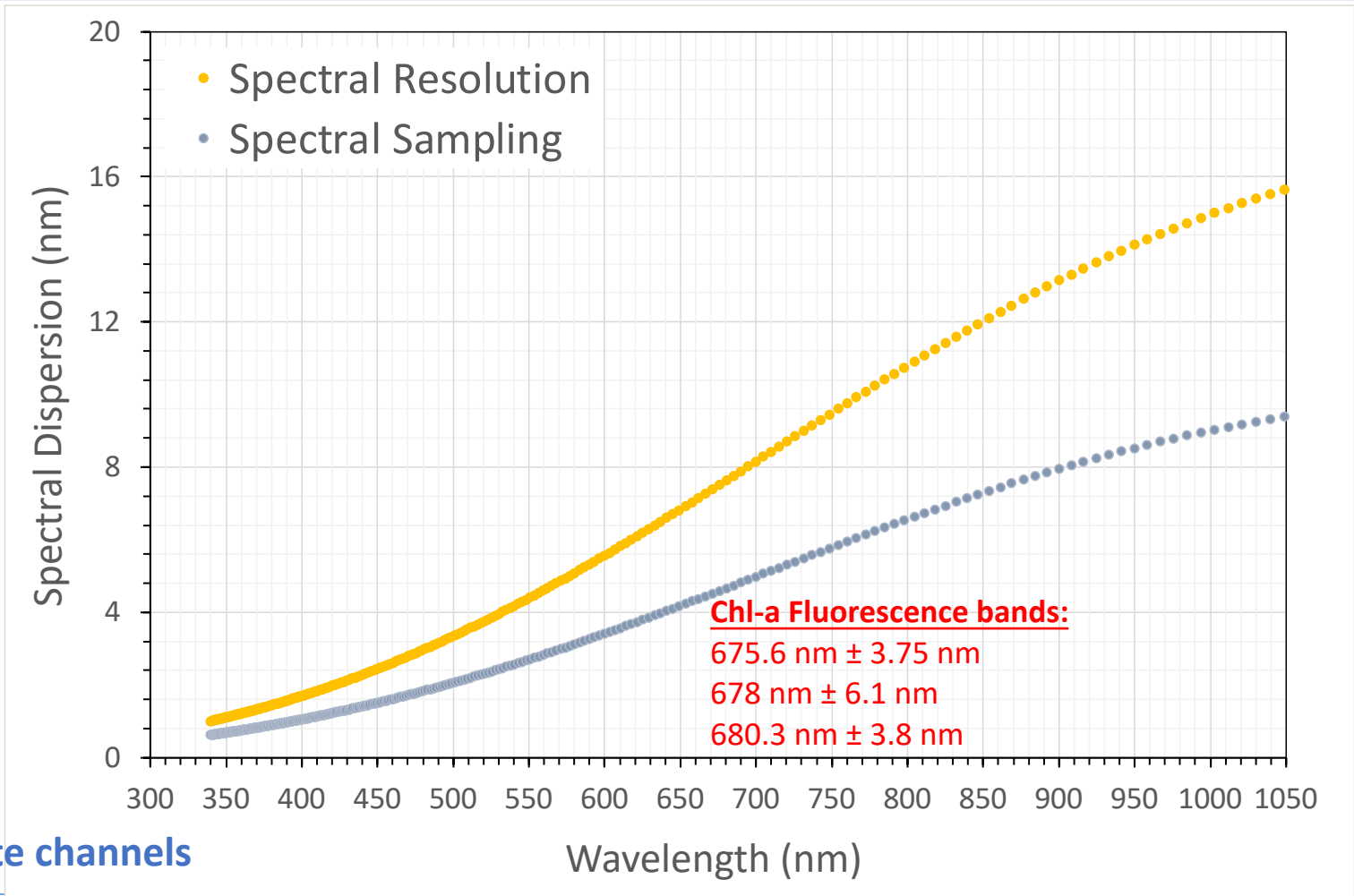


\*Anticipated PACE overlap



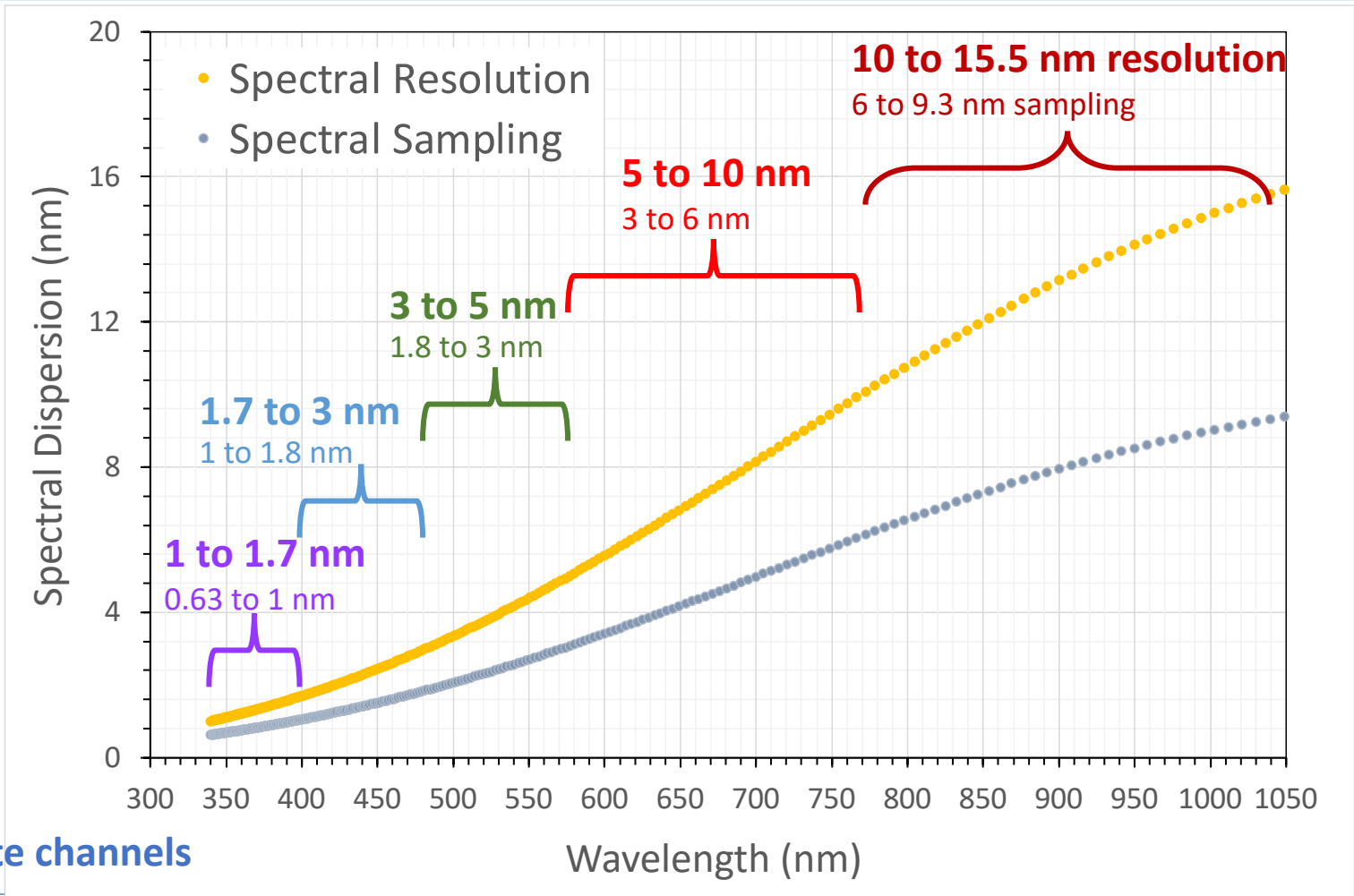
Schedule affords up to 5-7 daily matchup areas with PACE OCI year-round

# Modeled GLIMR Spectral Sampling and Resolution



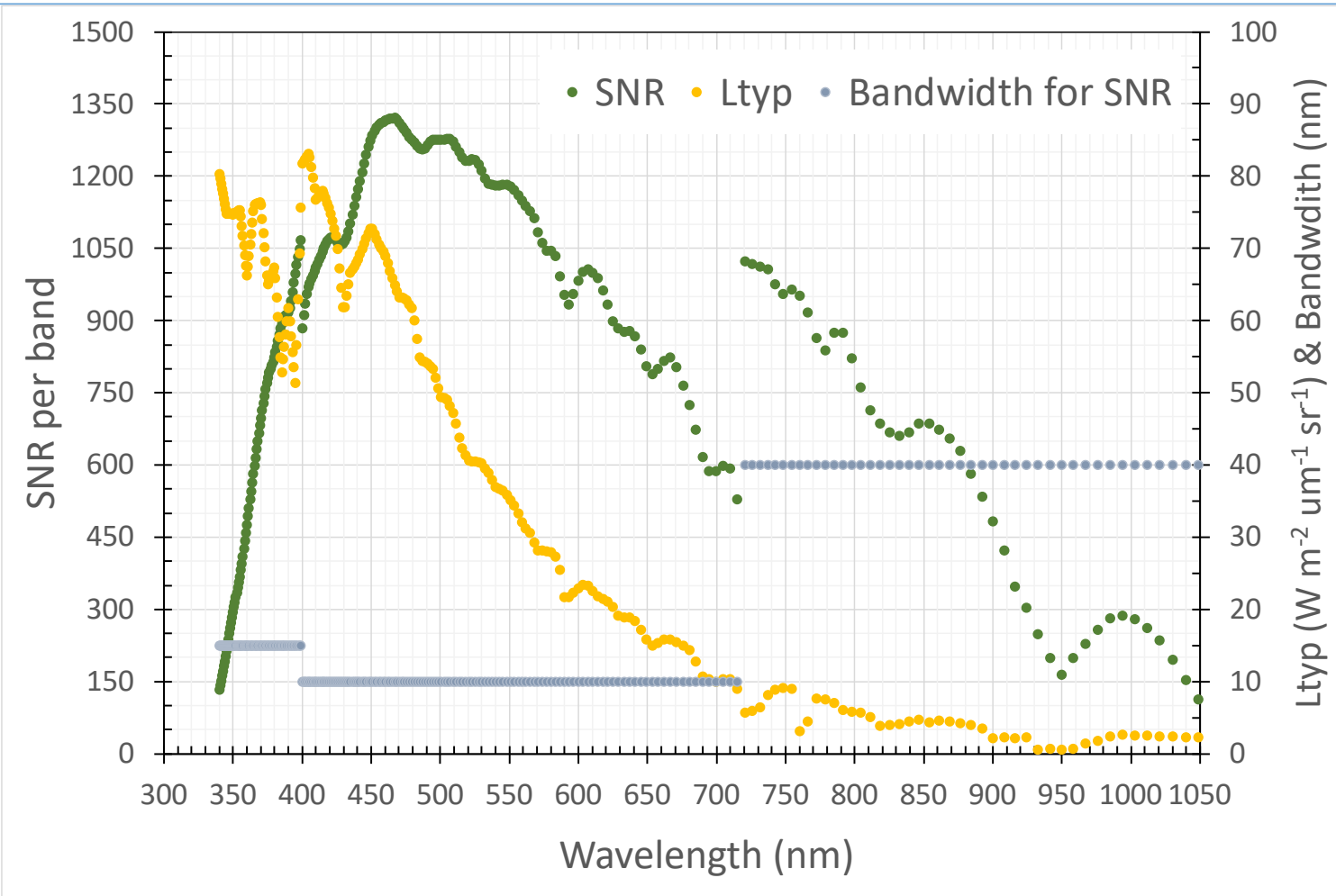
252+ discrete channels

# Modeled GLIMR Spectral Sampling and Resolution

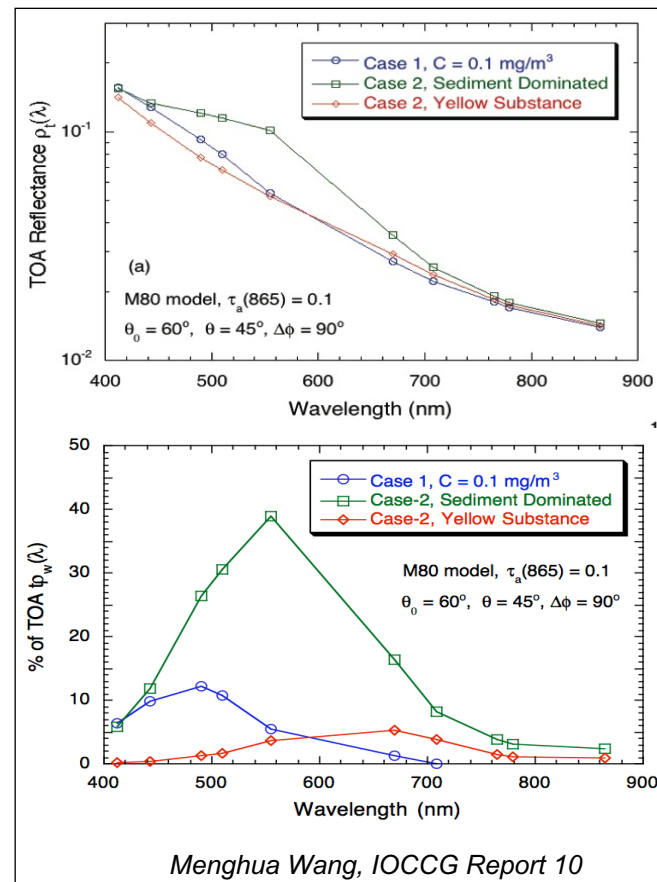


252+ discrete channels

# Modeled SNR as of June 2023



- Historically, the required performance goal for ocean color products (specifically, water-leaving radiance or reflectance) is 5% uncertainty.
  - Top-of-Atmosphere (TOA) radiance requirement ascribed to ocean color instruments is typically 0.5% uncertainty.
  - **Goal for GLIMR is to achieve ~0.5% uncertainty in TOA radiances in UV-Vis after vicarious calibration**



Small uncertainties at TOA have potentially large impacts on downstream products.

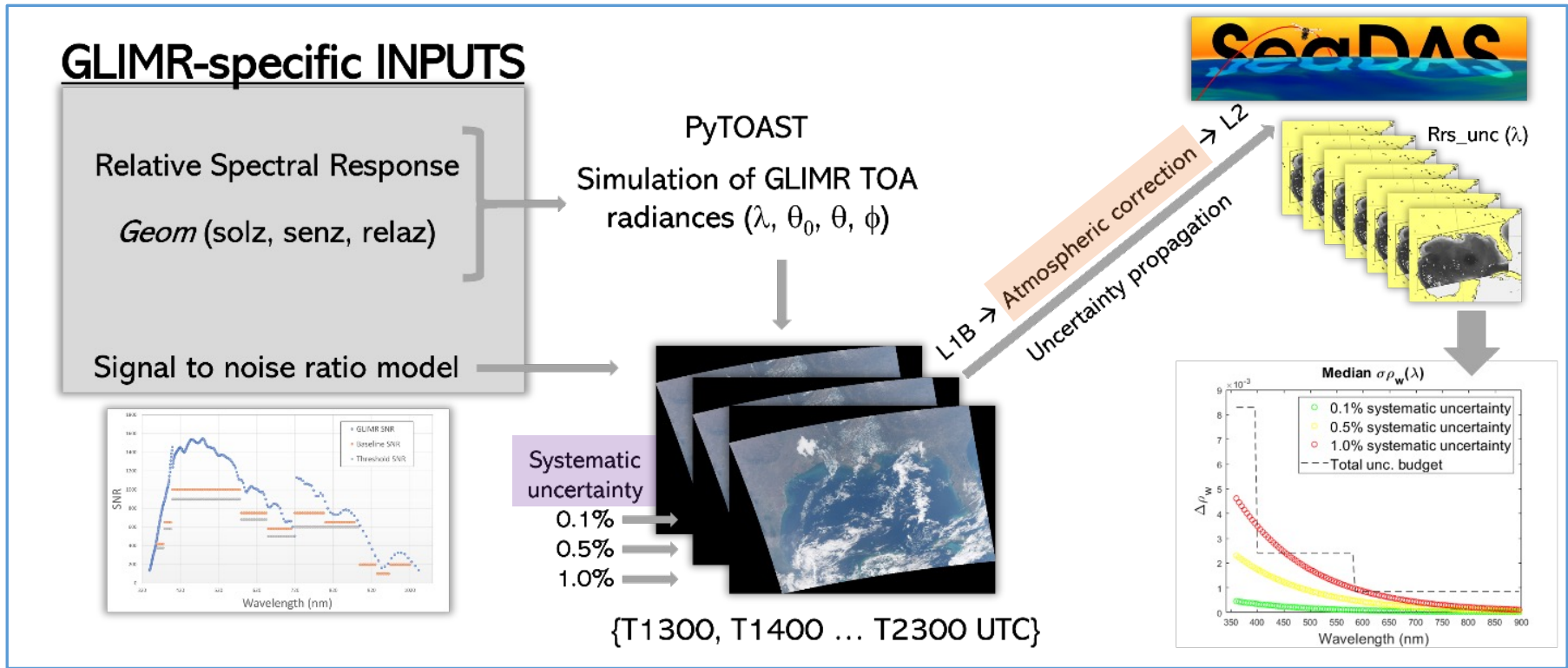
# Modeled TOA Radiometric Uncertainties – End-Of-Life

Ocean Color      Atmospheric Correction



	350-400	400-580	580-720	720-895	895-970	970-1040	Basis-of-Estimate
<b>Goal uncertainty (%) – 1 sigma</b>	0.5	0.5	0.5	0.5 to 1.13	3.0	2.0	Heritage ocean color for UV-Vis; NIR: from PACE OCI CBE + 20% as its EOL baseline
<b>Instrument On-Orbit Radiometric Uncertainty Estimate (%)</b>	0.485	0.438	0.518	1.12	3.02	1.93	Note: error terms summed by RSS for each header section (3 sections and at top level)
<b>Gain and Linearity Uncertainties</b>							RSS of K1, K2, K3, K5 and dn uncertainty terms; Radiometric stability, Temp., Linearity, dark counts
<b>K1: Absolute/Vicarious Gain</b>	0.20	0.10	0.10	0.696	2.11	1.67	Heritage; best option; based on PACE OCI uncertainty of vicarious calibration (340-720 nm) and absolute solar calibration (>720 nm)
<b>Image Artifact Uncertainties</b>							RSS of unc. from Stray light, high-contrasts, crosstalk, OOB, non-uniformity
<b>Polarization Sensitivity Residuals</b>							





**Systematic Uncertainty:** instrument + vicarious calibration uncertainties

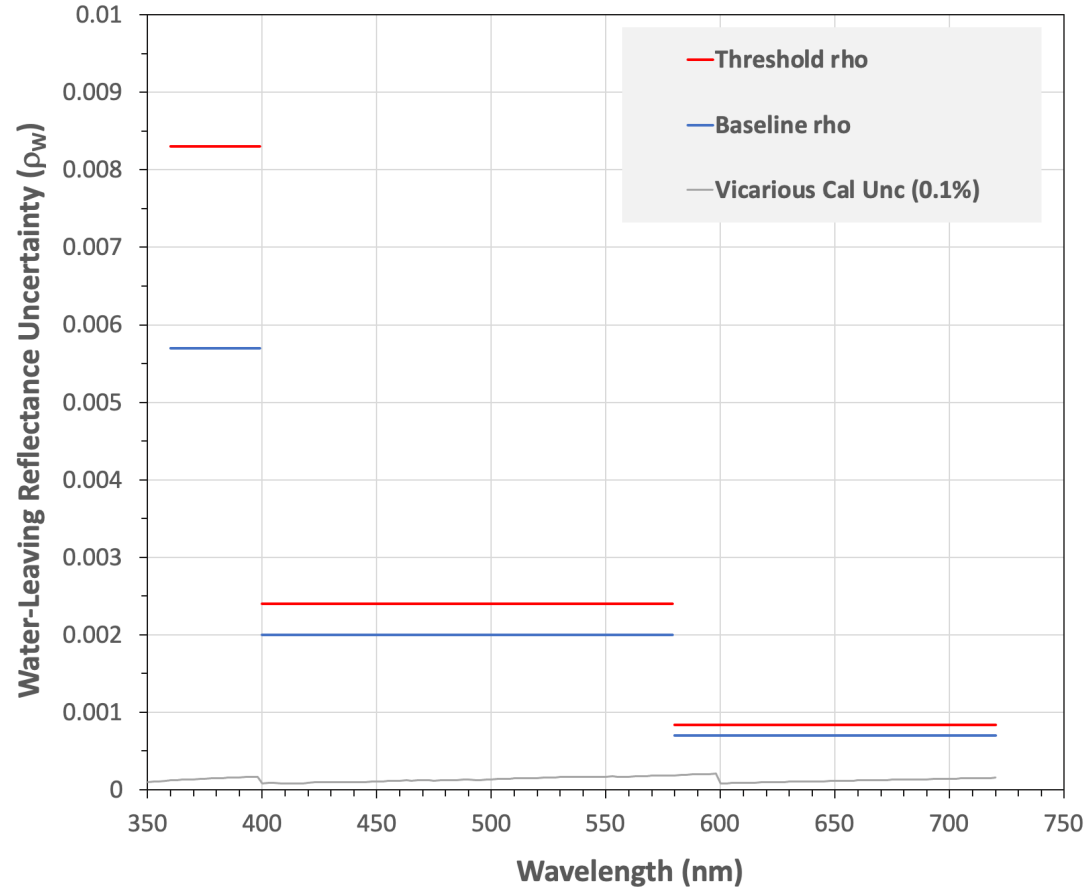
Science Data Product Modeling follows the approach implemented for PACE OCI

## Requirements

- Apply to bandwidths of 15 nm (or 10 nm over Fluorescence)

## Assumptions

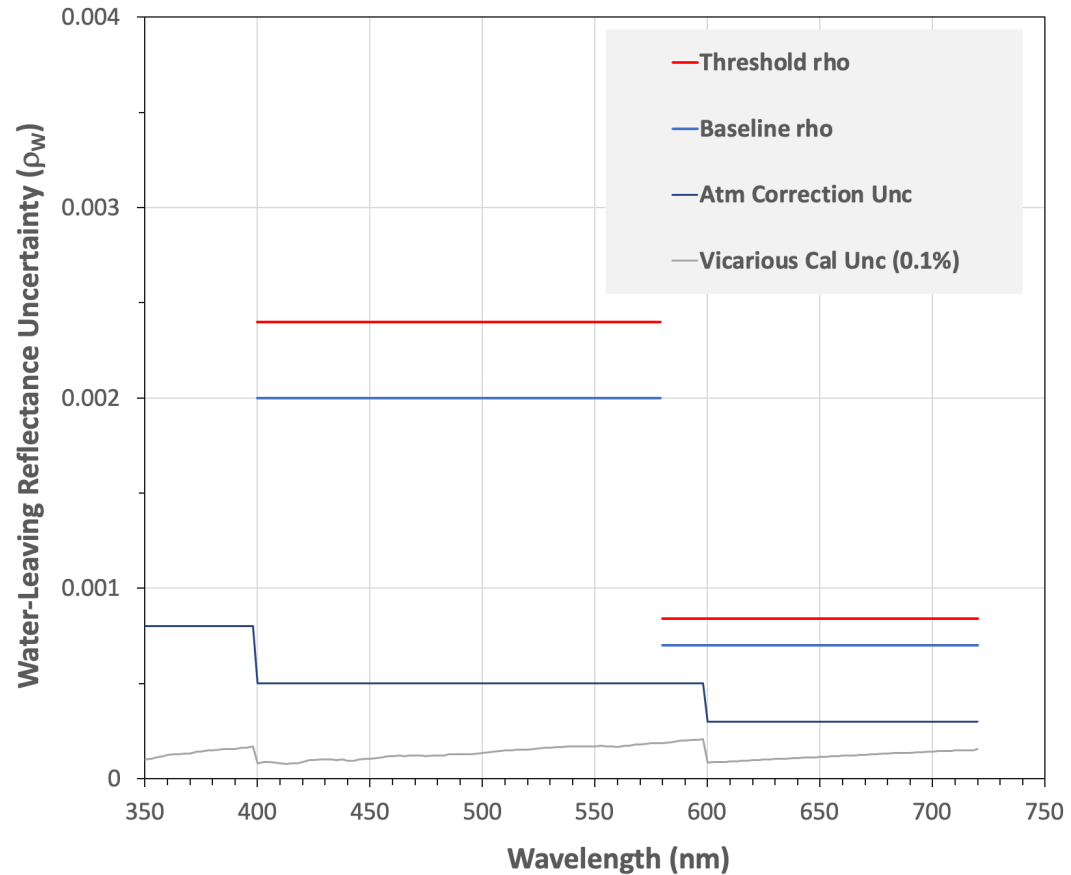
- Vic-Cal system(s) for PACE OCI adds 0.1% uncertainty
- Estimated at 0.2% for bands <400 nm



“rho” refers to  $\rho_w$  = water-leaving reflectance

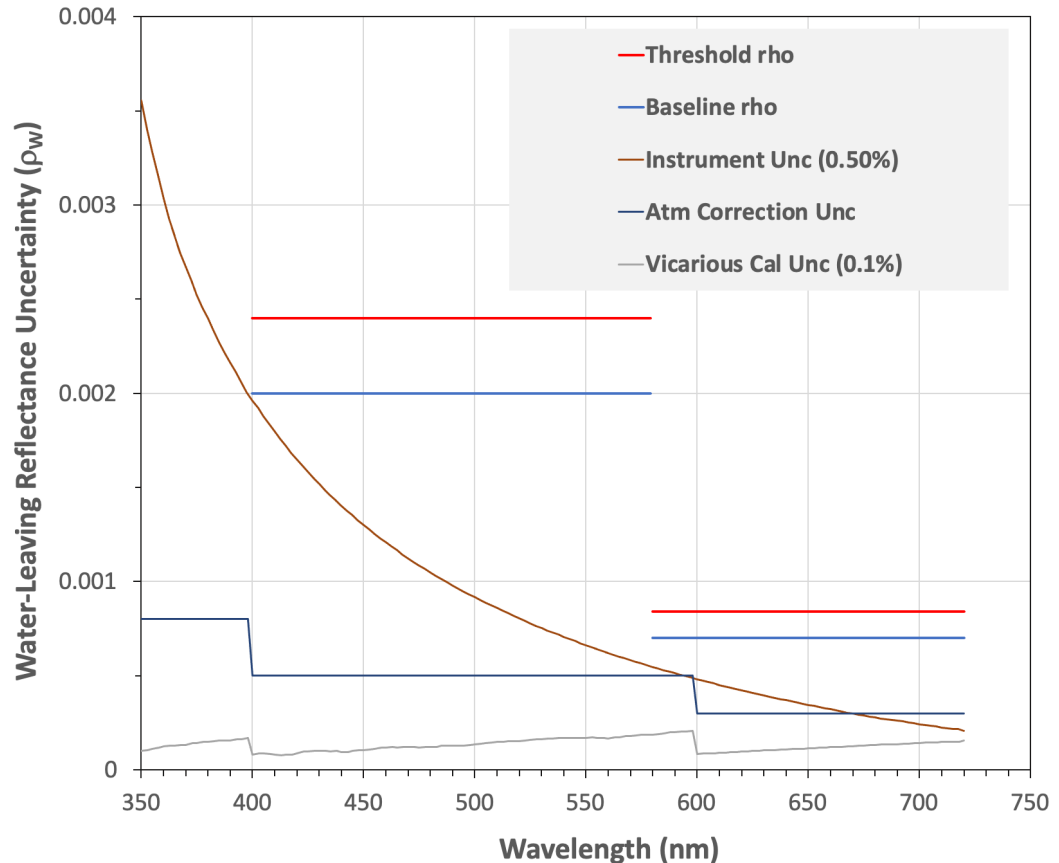
## Assumptions

- Applying current heritage Atmospheric Correction Algorithm (MSEPS: Multi-Scattering Epsilon)



## Assumptions

- Instrument uncertainty modeling follows PACE OCI (PyTOAST) but adapted to GLIMR (geometries)
- GLIMR CBE SNR
- 0.5% Radiometric Systematic Uncertainty attributed to instrument only (**Raytheon's precision requirement at EOL**)
- **Accounts for all instrument artifacts (T, polarization, stray light, non-linearity, crosstalk, drift, flat field uniformity, etc.)**
- Scenario for entire Gulf of Mexico

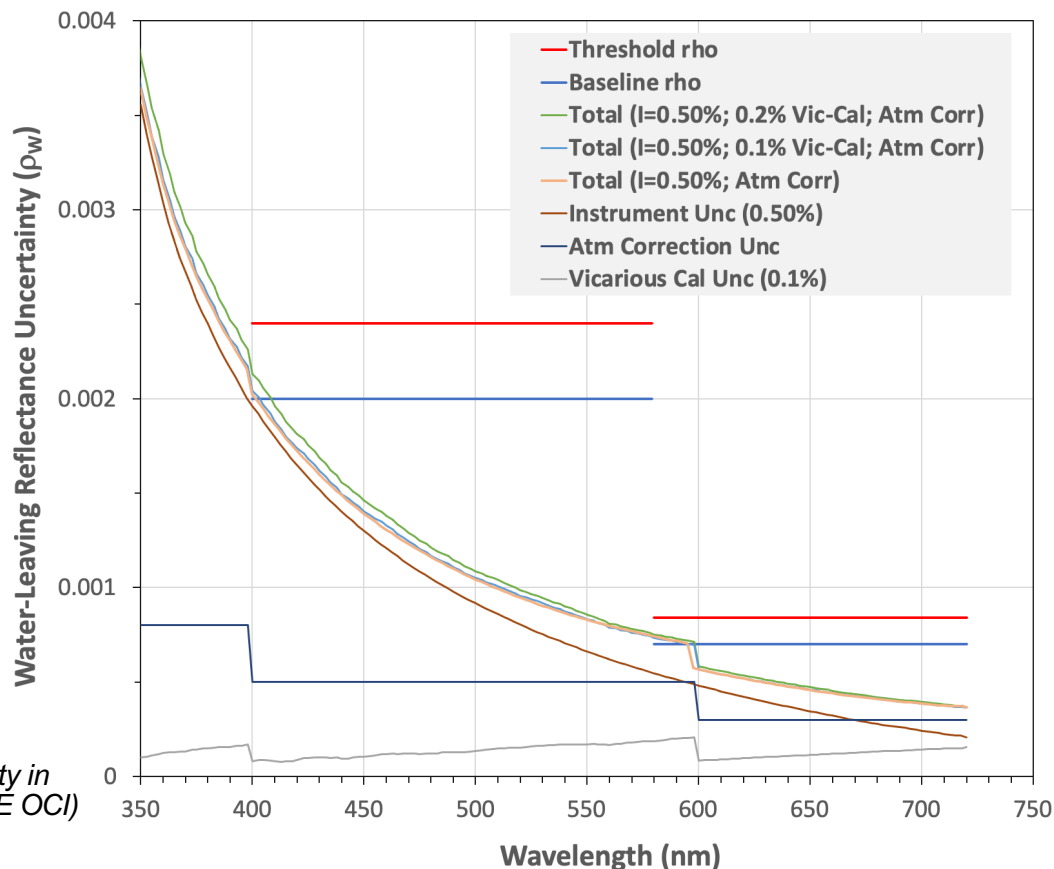


NOTE: (1) 0.69% Radiometric Systematic Uncertainty attributed for reference NIR band (870 nm) used to determine AOT as this band is not vicariously calibrated. (2) Approach follows PACE OCI's rigorously peer-reviewed (PRs, PDR, CDR, SIR, PSR) modeling and analysis

## Assumptions

- Instrument uncertainty modeling follows PACE OCI (PyTOAST) but adapted to GLIMR (geometries)
- GLIMR CBE SNR
- 0.5% Radiometric Systematic Uncertainty attributed to instrument (Raytheon's precision requirement at EOL)
- Added Vicarious Calibration & Atmospheric Correction to instrument on top of instrument uncertainty
- Scenario for deep ocean subset of Gulf of Mexico

No change in performance when model with 2% uncertainty in NIR reference band (from 0.69% relative value from PACE OCI)



**PLRA Threshold Requirements Fully Met and Baseline Met with Minor Exceptions**

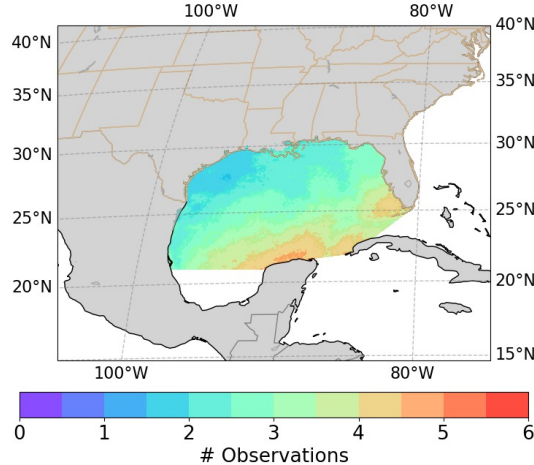
# Daily Cloud-Free Observations by Season

- GOES-East ABI Cloud mask from 2020.
  - ✓ Angular sampling distance: 56 microradia
  - ✓ Time resolution: 30 min
  - ✓ Subsatellite longitude 75W

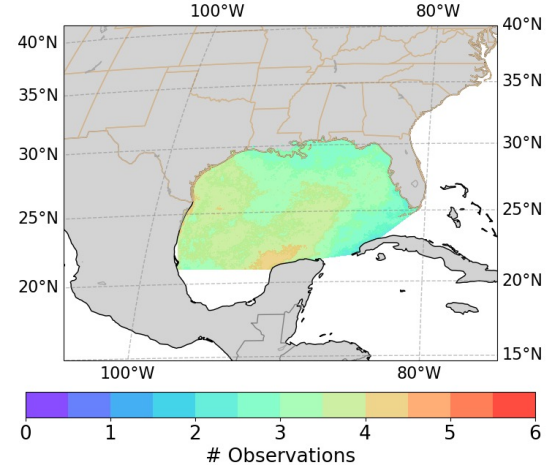
The number of cloud-free observations per day are averaged over each season.

by Boryana Efremova

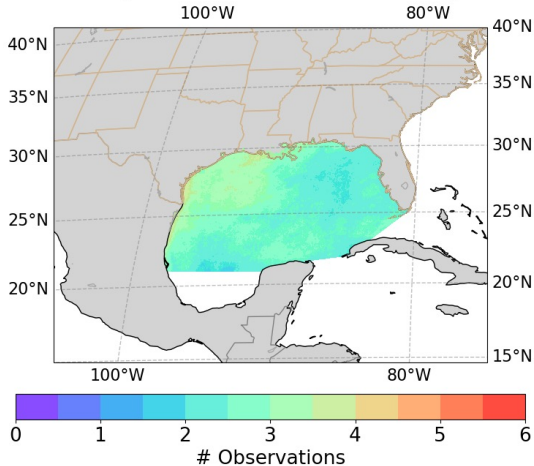
Daily Cloud-free Spring-average



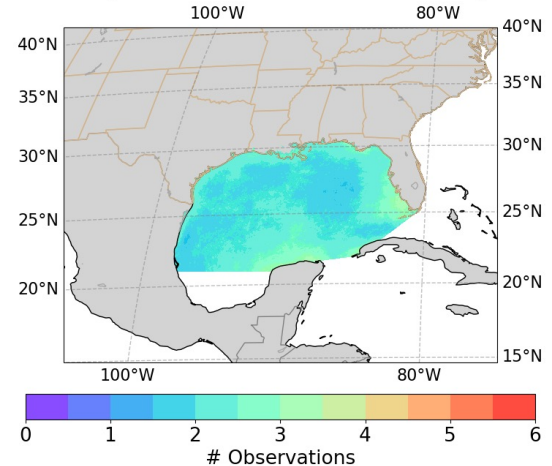
Daily Cloud-free Summer-average



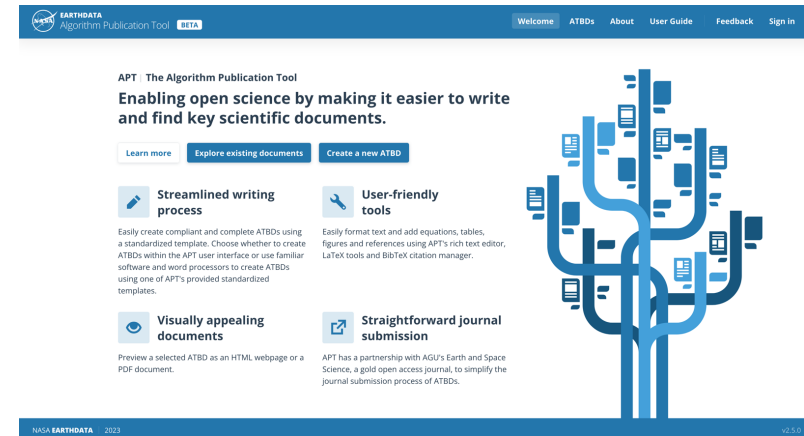
Daily Cloud-free Fall-average



Daily Cloud-free Winter-average



- APT (Algorithm Publication Tool)
  - APT is the "new" ATBD. While still in beta, it is available and anyone considering an algorithm should get an account and use APT to satisfy the ATBD requirement.
  - <https://www.earthdata.nasa.gov/apt/>



- SOT/SOB process (Science Operations Team/Board)
  - As the algorithms mature, we'll want the PIs to engage with the SOB/SOT process to get the algorithm into production.
  - The earlier the engagement the better.

# BACKUP



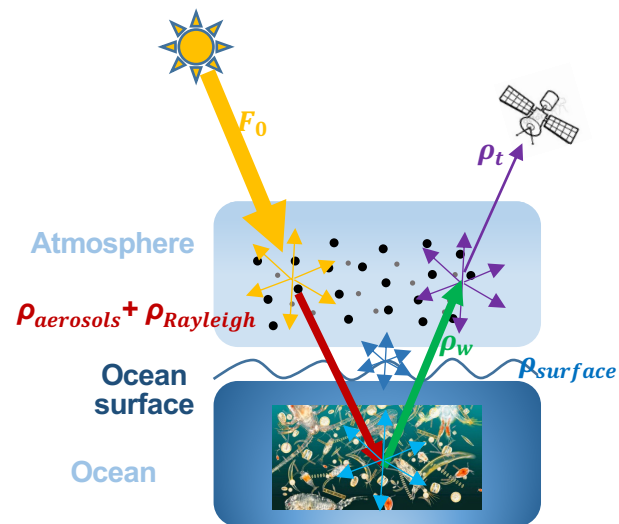
- Validation with data products from other satellite missions
  - PACE OCI, Sentinel 3A/3B/3C/3D OLCI, NOAA-20/21 VIIRS, etc.
- AERONET-OC (SeaPRISM) - [https://aeronet.gsfc.nasa.gov/new\\_web/ocean\\_color.html](https://aeronet.gsfc.nasa.gov/new_web/ocean_color.html)
- NASA SeaBASS - <https://seabass.gsfc.nasa.gov/> (ESDS-supported)
  - NASA and other field data collections
    - e.g., NASA supports collection of up to 3000 pigment samples annually
  - **Perform GLIMR validation matchups and uncertainty estimation**
- HYPERNETS - [http://www.hypernets.eu/from\\_cms/summary](http://www.hypernets.eu/from_cms/summary)
- Other federal/state/local government agencies; universities; etc.

VALIDATION APPROACH – Follow heritage and PACE methodologies

GLIMR must rely on other sources of in situ data as field validation efforts were descoped at KDP-C

Following methods for polarization sensitive instrument\*

- Daily matching of GLIMR pixels with OCI normalized water-leaving radiance ( $nL_w$ )
  - e.g., 3x3 OCI pixel bins ( $\sim 3.6 \times 3.6$  km)
  - Case 1 waters with depths greater than 1000 m; low AOT and Chl- $a$ ; homogeneous AOT and Chl- $a$
- Inverse processing of GLIMR (L2 to L1B) to bring OCI  $nL_w$  to the TOA and output modeled pixel Stokes vectors,  $[L_t, Q_t, U_t, 0]^T$ , at GLIMR wavelengths and GLIMR solar and viewing geometries.
  - Band radiances adjusted to match GLIMR bands
  - BRDF effects are accounted for in the propagation of the radiances from OCI to GLIMR viewing and path geometries
- Screen TOA pixels (following quality criteria) and generate datasets of radiance pairs  $[L_t, Q_t, U_t, 0]^T; L_m^G$ ) and ancillary information, detector element and time, geographic coordinates, solar and viewing geometries, and glint reflectance.
- Derive  $M_{11}$ ,  $M_{12}$ , and  $M_{13}$  per band and detector element.
- Compute ratio  $L_t / L_m^G$  to derive the cross-/vicarious gain coefficients (K1) per band

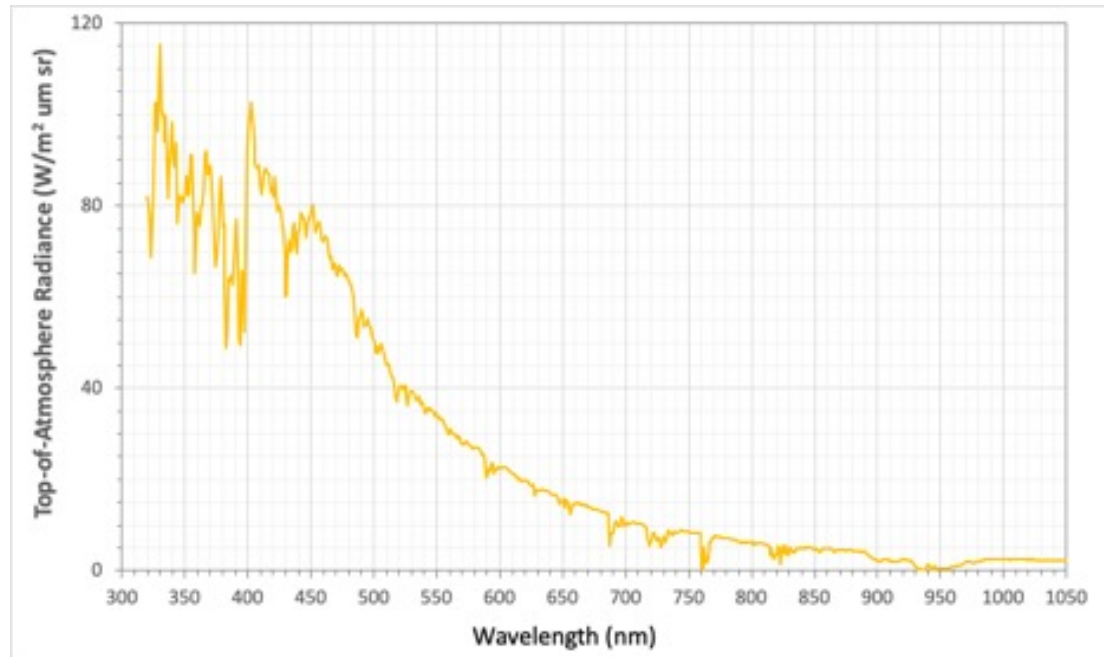


\* Kwiatkowska, Franz, Meister, McClain, & Xiong (2008). Cross calibration of ocean-color bands from Moderate Resolution Imaging Spectroradiometer on Terra platform," *Appl. Opt.* 47, 6796-6810.

$L_m^G$  = actual GLIMR TOA radiances

# Spectral Response Trending

- Potential change in dispersion over time on orbit. Use emission and absorption lines to track any changes



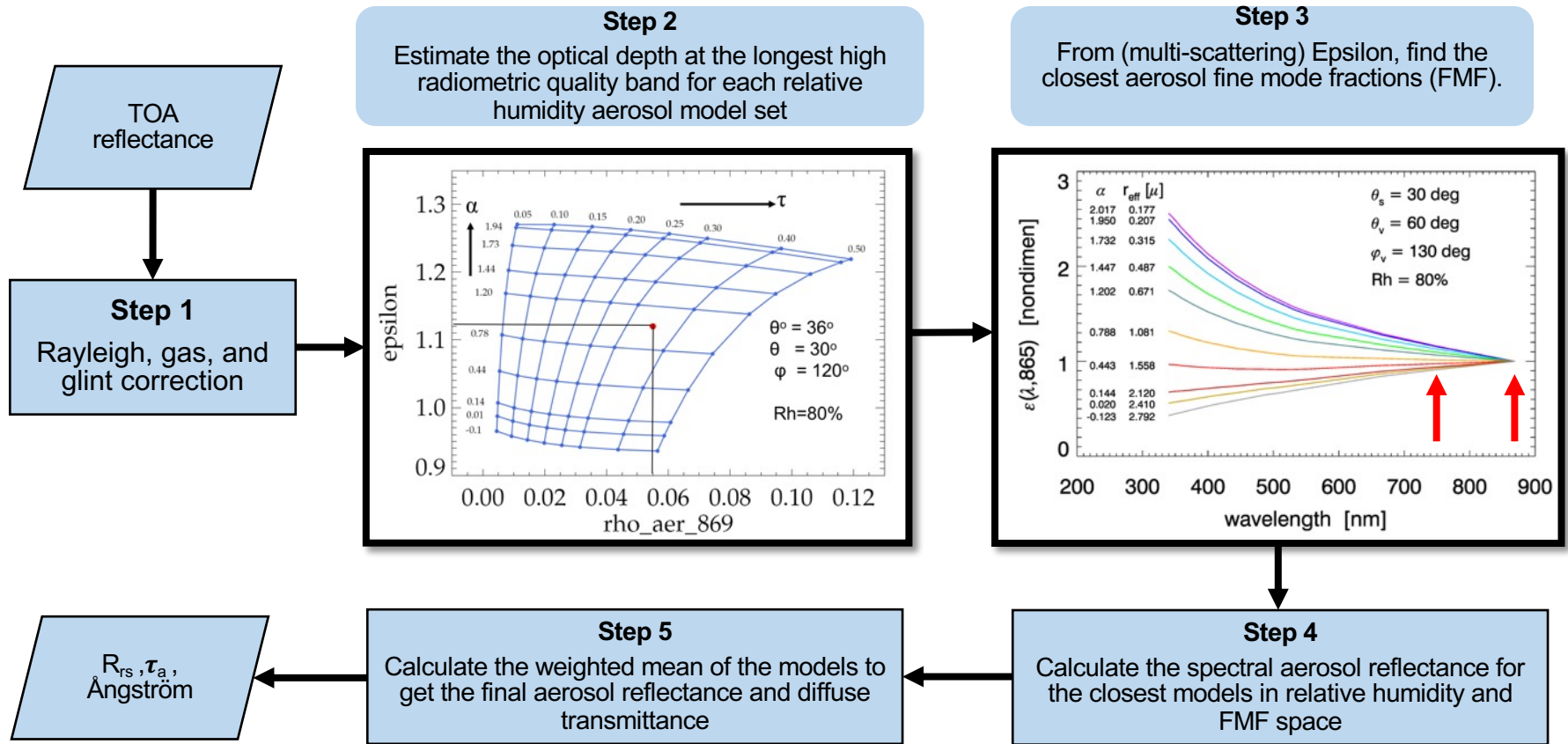
## 1. Fraunhofer solar emission lines:

- Will use at least two lines
- 434.048 nm (Hg), 588.997 nm (Na D<sub>2</sub>), 656.281 nm (Ha), 866.217 nm (Ca II)

## 2. Atmospheric absorption lines:

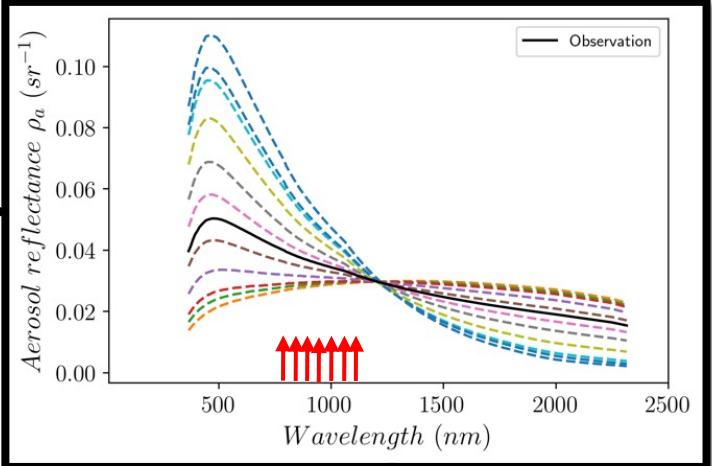
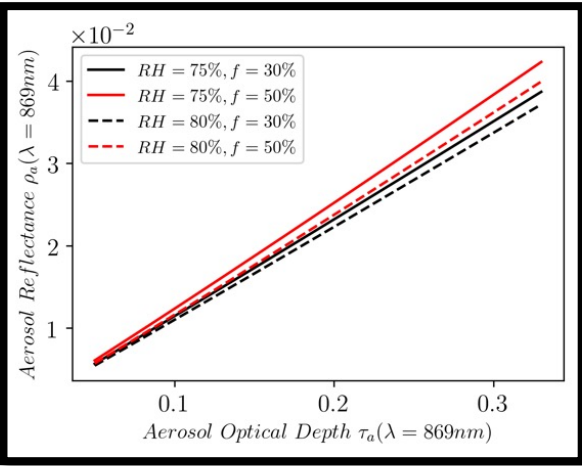
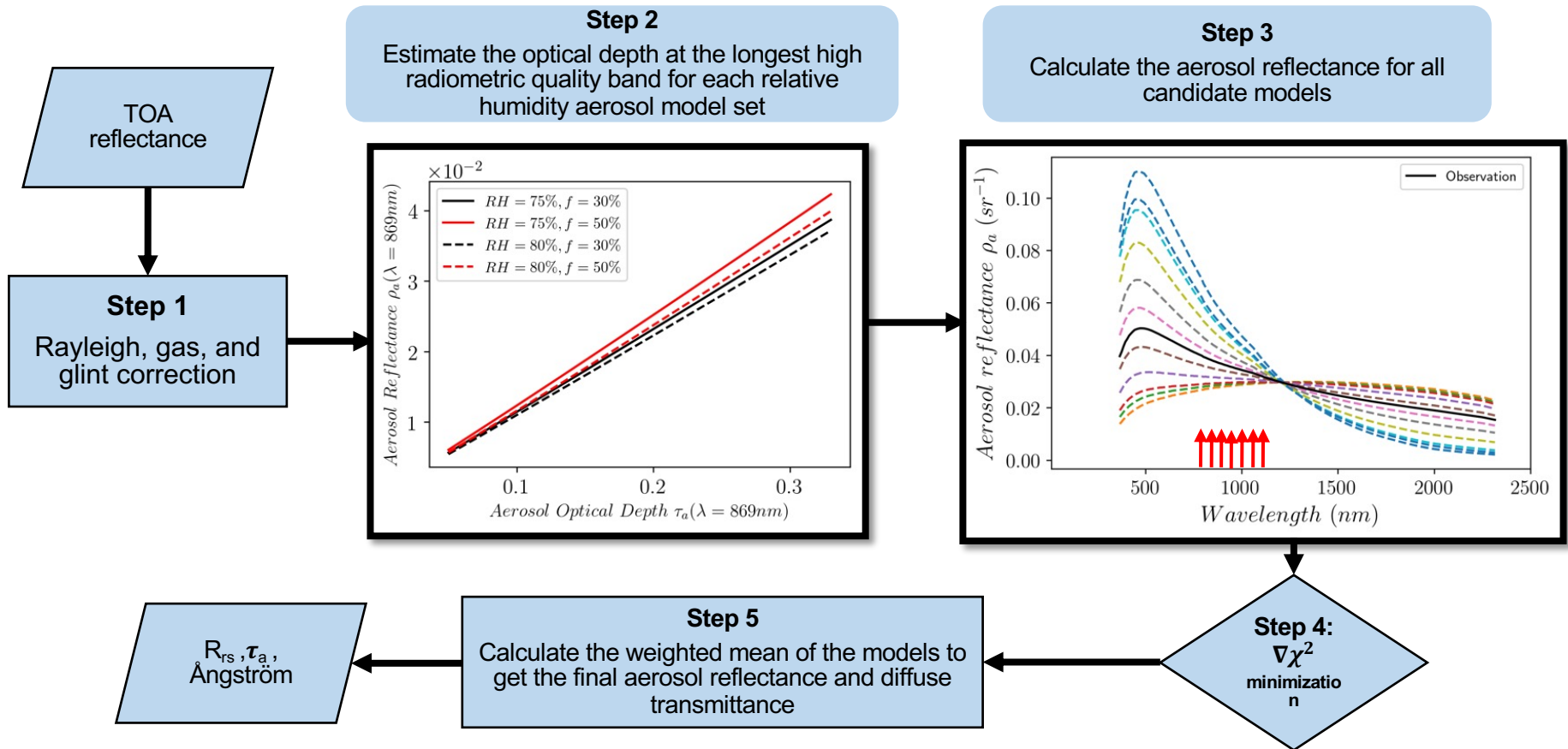
- 686.719 nm (O<sub>2</sub> B), 759.370 nm (O<sub>2</sub> A), 822.696 nm (O<sub>2</sub> Z), 898.765 nm (O<sub>2</sub> γ)

# Multi-Scattering Epsilon (MSEPS)



Ahmad, Z. and B. Franz (2018), Uncertainty in aerosol model characterization and its impact on ocean color retrievals, in PACE Technical Report Series, Volume 6: Data Product Requirements and Error Budgets (NASA/TM-2018 - 2018-219027/ Vol. 6)

# Multi-Band Atmospheric Correction (MBAC)



$$\chi^2(RH, f) = \frac{1}{DOF} \times \sum_{\lambda=\lambda_s}^{\lambda_t} \frac{[\rho_{obs}(\lambda) - \rho_a(\lambda, RH, f)]^2}{\sigma^2(\lambda)} \times SW(\lambda)$$

Ibrahim, A., Franz, B. A., Ahmad, Z., & Bailey, S. W. (2019). Multiband atmospheric correction algorithm for ocean color retrievals. *Frontiers in Earth Science*, 7, 116.

- extend MBAC to utilize UV spectral range in UV-dark waters
- include gas absorption in Rayleigh/aerosol LUTs to capture coupling
- include wind speed-dependent glint model in Rayleigh/aerosol LUTs
- improve correction for bi-directional reflectance (BRDF)

*Twardowski M, Tonizzo A. Ocean Color Analytical Model Explicitly Dependent on the Volume Scattering Function. Applied Sciences. 2018; 8(12):2684. <https://doi.org/10.3390/app8122684>.*

- implement ocean-atmosphere simultaneous retrieval (UV-VIS-NIR-SWIR)

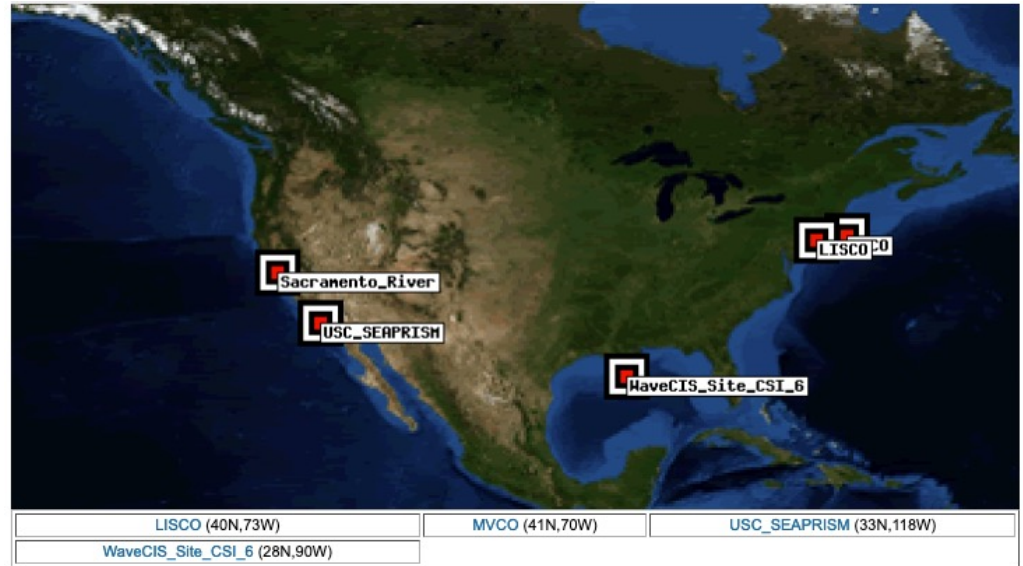
*Ibrahim, A., B.A. Franz, A.M. Sayer, K. Knobelspiesse, M. Zhang, S.W. Bailey, L.I.W. McKinna, M. Gao, and P. J. Werdell, "Optimal estimation framework for ocean color atmospheric correction and pixel-level uncertainty quantification," Appl. Opt. **61**, 6453-6475 (2022).*

# Approaches to Accomplish Vicarious Calibration (K1)

- Heritage MOBY approach - *in situ* instrumentation at appropriate field site(s) (e.g., Eplee et al. 2001; Franz et al. 2007; Zibordi & Melin 2017)
  - PACE vicarious calibration site or instrumentation (MarONet; HyperNAV)
- Alternate field site/instrumentation (e.g., Bailey et al. 2008; Zibordi et al. 2015)
  - AERONET-OC data; extrapolate spectrally
  - Hyperspectral sensors: in-water or above-water
    - HypSTAR in development at Tartu with ESA funding
    - Existing COTS with more extensive lab and field calibration
- Ocean surface reflectance model (Werdell et al. 2007)
- Ocean color satellite climatology of South Pacific waters (Franz et al. 2007; Concha et al. 2019)

Many options available to vicariously calibrate GLIMR

- AERONET and –OC are NASA-funded programs
- Bands: 400, 412.5, 442.5, 490, 510, 560, 620, 665, and 667 nm
- Additional bands at 709, 865, and 1020 nm for quality checks, turbid water flagging, and for the application of alternative above-water methods (Zibordi et al. 2002).
- Most useful sites for GLIMR
  - WaveCIS in Gulf of Mexico
  - USC off southern California
  - MVCO – Cape Cod



[https://aeronet.gsfc.nasa.gov/new\\_web/ocean\\_color.html](https://aeronet.gsfc.nasa.gov/new_web/ocean_color.html)



- Planned “new hyperspectral radiometer integrated in automated networks of water and land bidirectional reflectance measurements for satellite validation”
- Consortium of institutions coordinated by RBINS (Royal Belgian Institute for Natural Sciences)
- Funded by EU Horizon 2020

[http://www.hypernets.eu/from\\_cms/summary](http://www.hypernets.eu/from_cms/summary)

## WATER SITE - Rio de la Plata

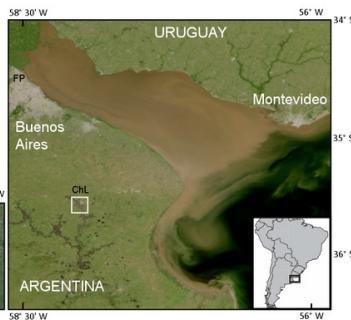


Fishermen pier in extremely turbid waters.

Coordinator: CONICET

Coordinates: 34.560865 S and 58.39881167 W.

## WATER SITE - Chascomus Lake



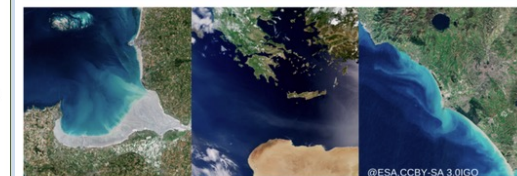
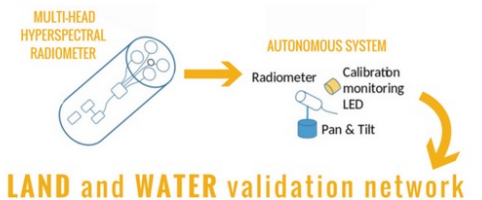
Chascomus Lake  
35° 35' 00" S - 58° 01' 10" W

Shallow extremely turbid inland water site.

Coordinator: CONICET

Coordinates: 35.58281326 S and 58.02024078 W

Embedding of the HYPERNETS sites into water and land surface reflectance validation networks based closely on the existing AERONET-OC and RadCalNet networks.



**Validation of surface reflectance at all spectral bands of all optical missions**

SENTINE-2, SENTINEL-3, LANDSAT, MODIS AQUA and TERRA, VIIRS, LANDMAPPER, PROBA-V, SABIAMAR, GOCI, GIGAT, AVHRR, SHALOM ENMAP, FLEX, GCOM, HYPKIM, PRISMA CHRIS, HIMAWARI, SKYSAT, GOES3, PACEROCI, PLANET-SCOPE, WORLDVIEW, NANOSATELLITES, ...

- **Radiometric Characterization and Calibration of Landsat and portable to GLIMR**
  - Derive surface spectral reflectance following established methodology
  
- **Pseudo Invariant Calibration Sites (PICS)**
  - Essentially invariant over time
  - Are spatially very uniform, have stable spectral responses over time
  - Atmospheric effects on upwelling radiance is minimal due to high surface reflectance
  - Are in regions where rainfall is extremely limited:
    - Prevents vegetative growth
    - Very sparse human populations

- TOA reflectance
  - at 30 min intervals from 9 am to 3 pm local time
  - Hyperspectral at 10 nm steps from 400 nm to 2500 nm

- Railroad Valley Playa
  - 15 x 15 km useable area
  - 38.50° N, 115.69° W
  - AERONET & RadCalNet



<https://www.radcalnet.org/#/>

Bouvet et al. 2019 Rem Sens; <https://doi.org/10.3390/rs11202401>

RadCalNet  
NASA Goddard Space Flight Center

CEOS Committee on Earth Observation Satellites

Railroad Valley Playa  
return to site description

Data

Monthly Aerosol Optical Depth at 550nm (Month 05)

Webcam views for the day

Download daily data for all instruments

Sensor : 00    Input version : 05    Output version : 04

Atmospheric parameters    BOA Reflectance    TOA Reflectance

ALL

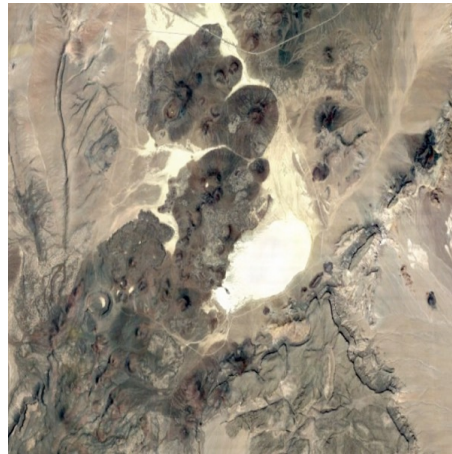
TOA Reflectance

Wavelength (nm)

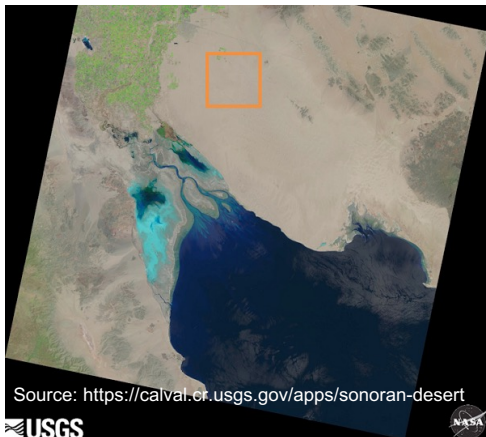
Legend for TOA Reflectance plot:

- 17:00
- 17:30
- 18:00
- 18:30
- 19:00
- 19:30
- 20:00
- 20:30
- 21:00
- 21:30
- 22:00
- 22:30
- 23:00

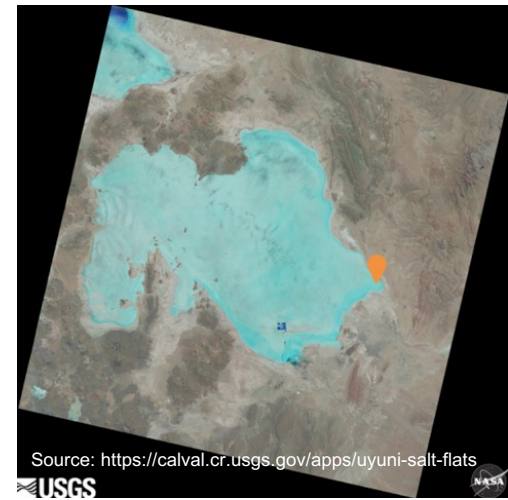
- Landsat; AERONET systems
- Lunar Lake Playa, Nevada
  - Dry lake bed; <0.5% reflectance variance
  - 1.5 x 2.5 km usable area
  - 38.4° N, 115.99° W
- Sonoran Desert, Mexico
  - spatially uniform
  - 15 x 15 km usable area
  - 32.35° N, 114.65° W



- Uyuni Salt Flats, Potosi, Bolivia
  - 25 x 25 km usable area
  - 20.38° N, 66.95° W



Source: <https://calval.cr.usgs.gov/apps/sonoran-desert>



Source: <https://calval.cr.usgs.gov/apps/uyuni-salt-flats>